IBIS/ISGRI STATUS

Philippe Laurent, Jérome Rodriguez, Floriane Cangemi, Christian Gouiffès, Aleksandra Gros, Aymeric Sauvageon, Geoffrey Daniel, Alexis Coleiro, Sarah Antier

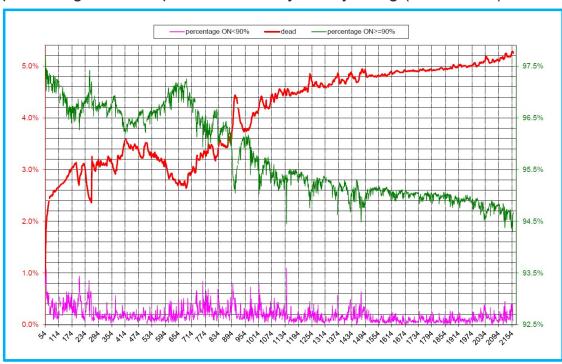
CEA/DRF/IRFU/DAp & APC

IUG#23 26/11/2019

ISGRI OPERATION

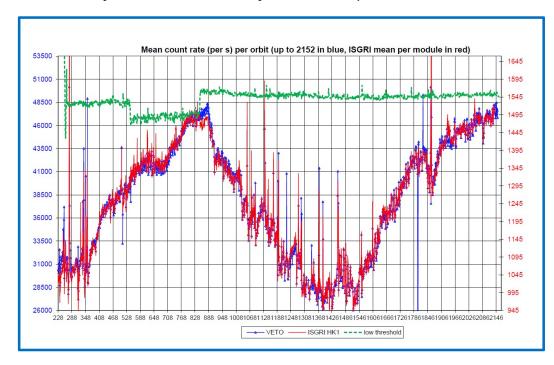
Dead pixels

The percentage of dead pixels is still very slowly rising (around 5%).



ISGRI/Veto Background

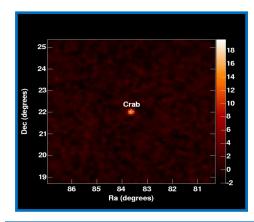
As solar activity goes to minimum, ISGRI raw count rate increases (due to cosmic rays increase, seen by VETO rate).

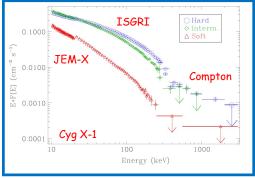


COMPTON SOFTWARE

Compton mode data analysis

- Compton: Delivery to ISDC of Compton software to make images and spectra is underway.
- We made a first delivery of ic_Mosalma and dal3ibis-calib to ISDC.
- Dal3ibis-calib is OK.
- Some works remain to be done at CEA for ic Mosalma.
- We plan to write a paper on the Compton mode software development in the next months.
- OSA11: We have begun to write the frame of the ISGRI calibration report.





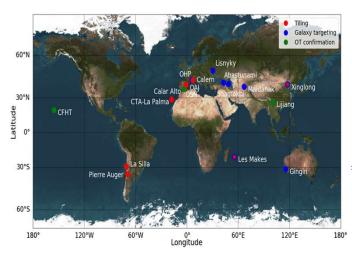
lovember 26th 2019 **7**

SCIENCE

Multi-messengers, polarization, compact objects

- On-going multi-messenger activity at Paris/Saclay
 - Participation to the MM team shift (A. Coleiro).
 - > Development of a true Compton imaging software (G. Daniel).
 - Development of the "GrandMa" network (S. Antier).
- Several polarization studies are on-going :
 - Cygnus X-1 (Cangémi, F. & Rodriguez, J.)
 - > V404 Cygni (Chang, B. PHD thesis)
 - > GRBs (Götz, D.)
 - Crab total and phase-resolved (Gouiffès, C.)
- CDD 1 year CNES position for science studies in discussion ...

Global Advanced Network Devoted to Multi-messenger Addicts



- Coordination of the telescopes
- 27/33 alerts distributed for Ö3a
- 16 min for minimal reactivity

Antier et al., MNRAS, 2019

23 telescopes - 17 observatories 11 countries - 70 scientists

Identify and characterize the GW electromagnetic counterpart:

photo: up to 23 mag

spectro: up to 20 mag

V404 polarization (B. Shang)

- ➤ B. Shang PHD thesis partly on V404 Cygni held on November 5th in Hsinchu (Taïwan).
- Following the study shown during the Amsterdam conference (2016, see below), we decided to look at the variation of the polarization during rev. 1555 on scw timescale (change in PA and PF).

Revolution number	1554	1555	1556	1557	1558	1559	1560	1561	1562	1563
Exposure time (ks)	170	149	84	153	163	171	93	165	174	117
S/N	8.8	14	4.6	28	-0.4	-0.2	-0.05	1.26	1.75	1.09
Polarization detected (yes/no)	No	Yes	No	No	No	No	No	No	No	No

V404 polarization (B. Shang)

- Study of the variation of the polarization during rev.
 1555 on scw timescale (change in PA and PF).
- Study of the possible correlation with the presence/absence of magnetic field, measured through spectral fitting.





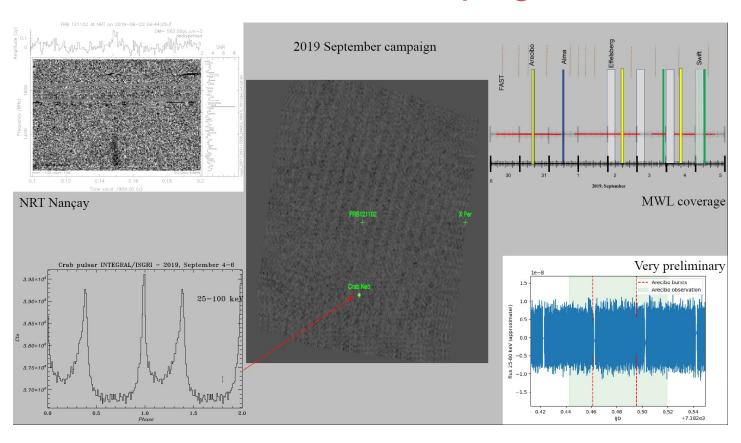
Context

- · Fast radio bursts (FRB) are millisecond-duration transient radio signals
- Discovered in 2007 from Parkes archival pulsar data survey (Lorimer's burst)
- From DM (Dispersion measure), extragalactic origin favored
- On 2019, Nov 100 FRBs registered (source : FRBCAT catalog), among them 10 repeaters
- Host galaxy identified in 4 cases
- · Fast and active magnetar inside a dense environment favored model but many alternative theories exist
- Several models predict extended gamma-ray emission
- Several multi-lambda programs

<u>History: September 2017</u>: Integral regular observations (2 revolutions), severa Radio facilities involved (Nançay, Effelsberg, Arecibo, Green Bank, Fast), Optical: OHP T193cm+GASP (Galway Fast polarimeter), T120cm+CCD, VHE participation (Hess, Veritas, Magic)

- → No radio bursts detected during this campaign (+ unforseen pb with Arecibo and NRT)
- → <u>New strategy</u> adopted for **2019 (AO16)** : <u>ToO mode</u>, trigger based on the regular radio monitoring of FRB121102 with NRT
 - → Source active in late August 2019 (NRT-Nançay) → ToO INTEGRAL (500 ksec)

2019 FRB Campaign



2019 FRB Campaign



FAST Detects Multiple Bursts from Fast Radio Burst 121102

Sep 09, 2019

Five-hundred-meter Aperture Spherical Radio Telescope (FAST), the world largest single dish radio telescope, began commissioning on 25th September, 2016. In February 2019, it announced call for proposal publicly to Chinese astronomers. More than 133 proposals from 21 institutions including the University of Hong Kong were received. Granted proposals started to arrange observations on 18th April, 2019.

In its recent tracking observation of the Fast Radio Burst (FRB) FRB121102 using FAST L-band 19-beam receiver (with FWHM of ~2.95' for individual beam), FAST detected multiple bursts.

The bursts were firstly identified by the FAST FRB backend on August 29th (UT), which performs real time signal processing of 19-beam data and automatic candidate selection/triggering. The subsequent single pulse search using multiple pipelines have turned up many tens of pulses with significant SNR in observations carried out so far, from 29th August to 3rd September (UT).

The total number of bursts detected from FRB121102 this time is known to be the highest by far. Careful cross-check and further processing are being carried out.

FAST FRB backend was developed by researchers from the National Astronomical Observatories of the Chinese Academy of Sciences (NAOC), and the collaborators from UC Berkeley, Beijing Normal University and Xinjiang Astronomical Observatory of the Chinese Academy of Sciences. It passed technical review and final project review in November 2018 and was rated as excellent NAOC project.

The FAST FRB backend system has high-efficiency real-time pulse capture capability, and can observe in parallel with most observation tasks. It will play an important role in the discovery of new FRBs, improving the position accuracy and capturing the high-resolution absorption lines generated by FRB in real time.

FRB are the brightest burst in radio-band currently known in the universe, but there is no reasonable explanation for their origin. The repeated bursts detection of FAST will promote the research on understanding the origin and physical mechanism of FRB.

FAST has been targeting FRB 121102 since April 2019, and is executing more observations under the auspice of engineering testing time and multiple approved PI-led programs. In addition to the regular on-going FRB follow-up programs, the current observation was also motivated by timely and valuable alerts from the colleagues in the JNTEGRAL team, Arecibo team, Max-Plank Institute for Radio Astronomy, Berkeley, and Cornell University.

| INTEGRAL and rated join joint programme of FRB 121102 during a renewed activity | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 1919 | 19

Address: http://english.cas.cn/ Copyright © 2003 - 2015 Chinese Academy of Sciences

New campaign in 2020 (AO17) in ToO, 3 orbits allocated