

IBIS Status Report:
Crab long term variability and
Matrices production



IUG 22-23
January 2013

P. Ubertinil L. Natalucci & IBIS team



Will INTEGRAL be the last to
retire?!...



1 mm W Hopper to limit
the CdTe F.o.V.

Detection system
128x128 CdTe detectors
Anticoincidence system
32 BGO slabs, 20 mm thick
Read out by 64 PMs
Energy range 0.175-10 MeV



ASI Contacts: Scientific & Technical Staff Status per January 2013

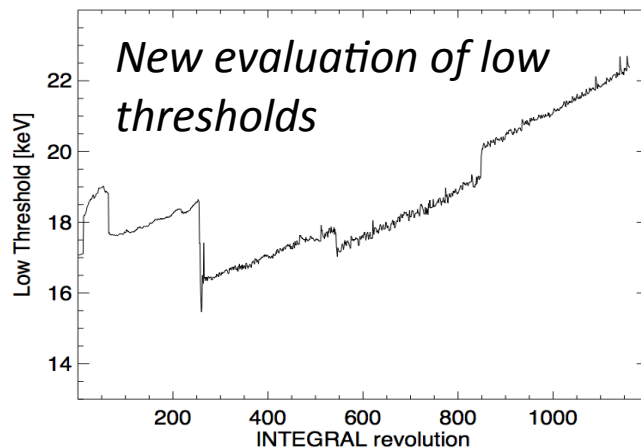
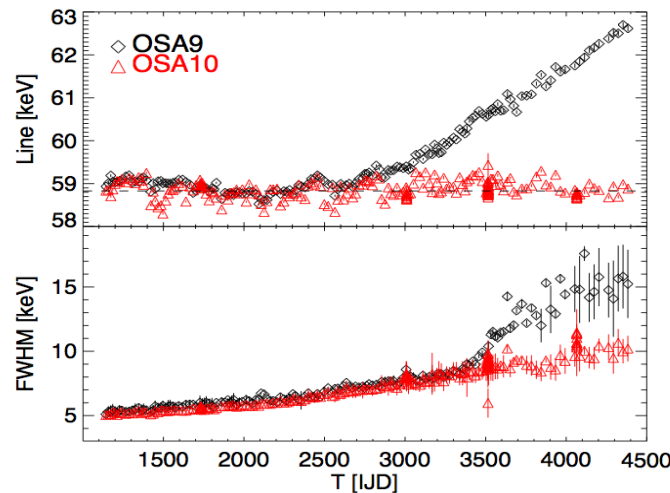
- A progress meeting has been held **on January 18** between the INTEGRAL program Italian coordinator (PU) and the ASI contract officer (EC).
- **ASI has confirmed the provision in the 2013-2016** planned budget of the financial support for the in flight INTEGRAL operation and calibration support.
- The final negotiation will be carried out at the end of the present 3 years contacts, formally expiring next **August 2013. The new contract is expected to start just after (August 2013).**

People on contract and fellowship FULL TIME now:

Contracts	Institute	Started	End Phase III
•M. T. Fiocchi	Art. 15 IAPS-Roma	1/10/2010	18/12/2012*
•R. Landi	Art. 15 IASF-Bologna	1/10/2010	31/08/2013
•F. Panessa	Art. 15 IAPS-Roma	1/10/2010	31/05/2013
•A. Tarana	Art. 15 IAPS-Roma	1/10/2010	30/06/2012
•Fellowship			
•M. Molina	Ass. Ric. IASF-Bologna	1 y	30/10/2013
•P. Parisi	Ass. Ric. IAPS	1 y	30/11/2013
•S. Lotti	Ass. Ric. IAPS	1 y	30/05/2013

IBIS/OSA10 calibration

- ◆ New OSA10 calibration of gain and energy correction, corrects many spurious effects left in OSA9
- ◆ 12 new ARFs delivered for the OSA10 distribution



PROCEEDINGS
OF SCIENCE

INTEGRAL IBIS/ISGRI energy calibration in OSA 10

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All details in the paper submitted as
POS (INTEGRAL Conf 2012)

Current response status/performance

- ◆ Check status using the most recent Crab data (revs. 1214-1221)
- ◆ One on-axis, plus 2 off-axis (mask) observations. Off axis angles range ~6-9.5 deg
- ◆ Folded by current response
- ◆ Count rate spectra not well modeled, also a feature near ~55 keV appears.
- ◆ Off axis data show departure from model especially at $E < 30$ keV

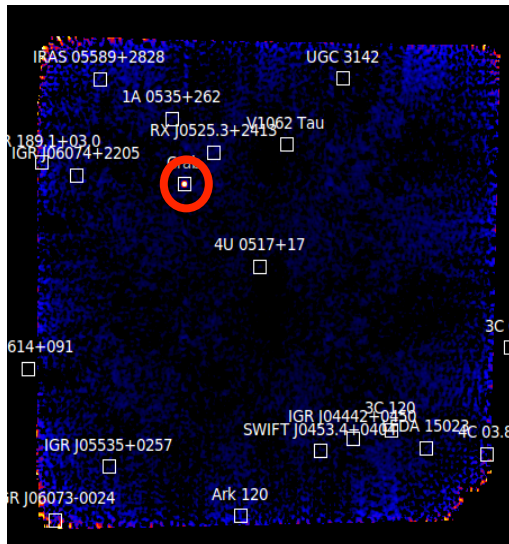
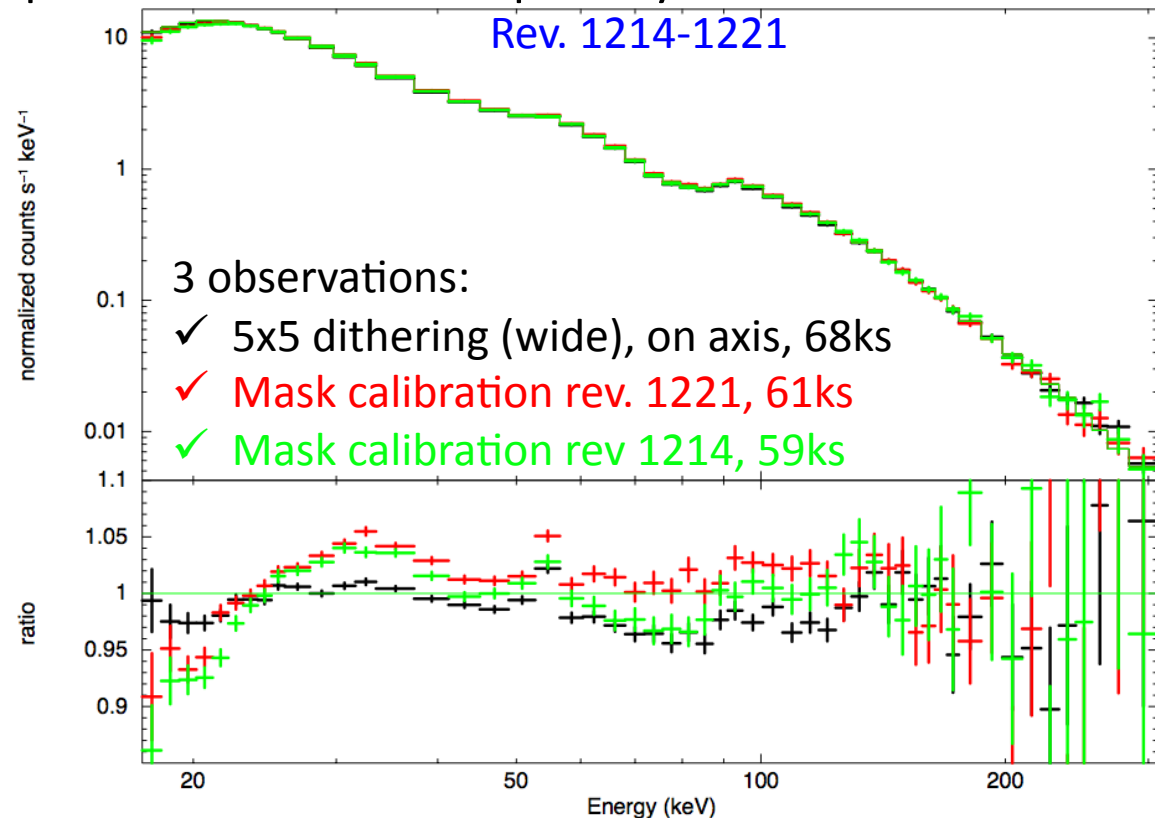
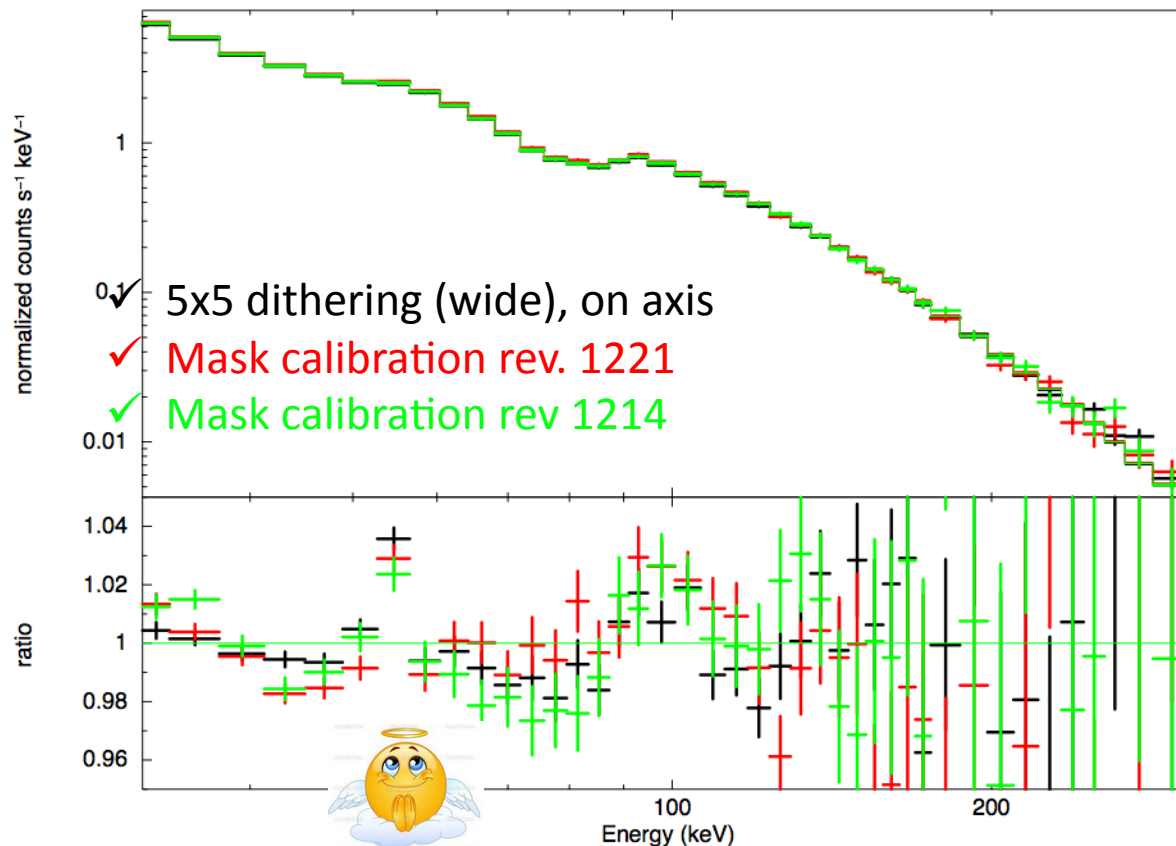


Image mosaic for mask calibration, rev. 1221 (similar for rev. 1214)



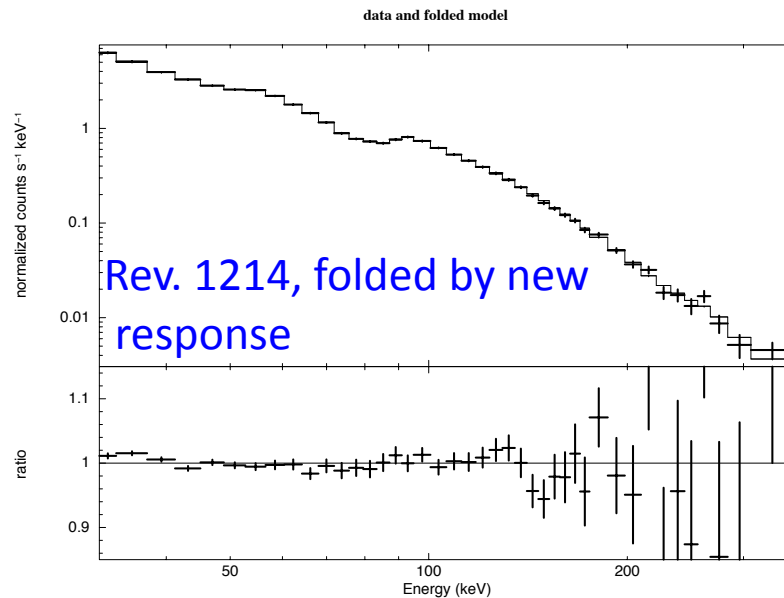
Fit > 30 keV, on- vs off-axis

- ◆ E>30 keV, fit with free normalization constants
- ◆ Yields a factor $\sim 2-3.5\%$ higher for the off-axis data respect to the on-axis data
- ◆ The fit is formally satisfactory ($\chi^2=1.06$ for syst=1.5%) but residual features appear especially at ~ 55 keV



New ARF for OSA10, rev. 1221

- ◆ The above tests show need for a new ARF
- ◆ We have computed one more ARF to account for the most recent observations. Using Crab dithering data of rev. 1221
- ◆ The ARF will be delivered to Saclay and tested
- ◆ We propose to re-calibrate ARF 2 times/year instead of 1/year. Use of all Crab periodic calibration observations
- ◆ The re-calibration procedure is standard, and fully supported.



Fit with broken power law

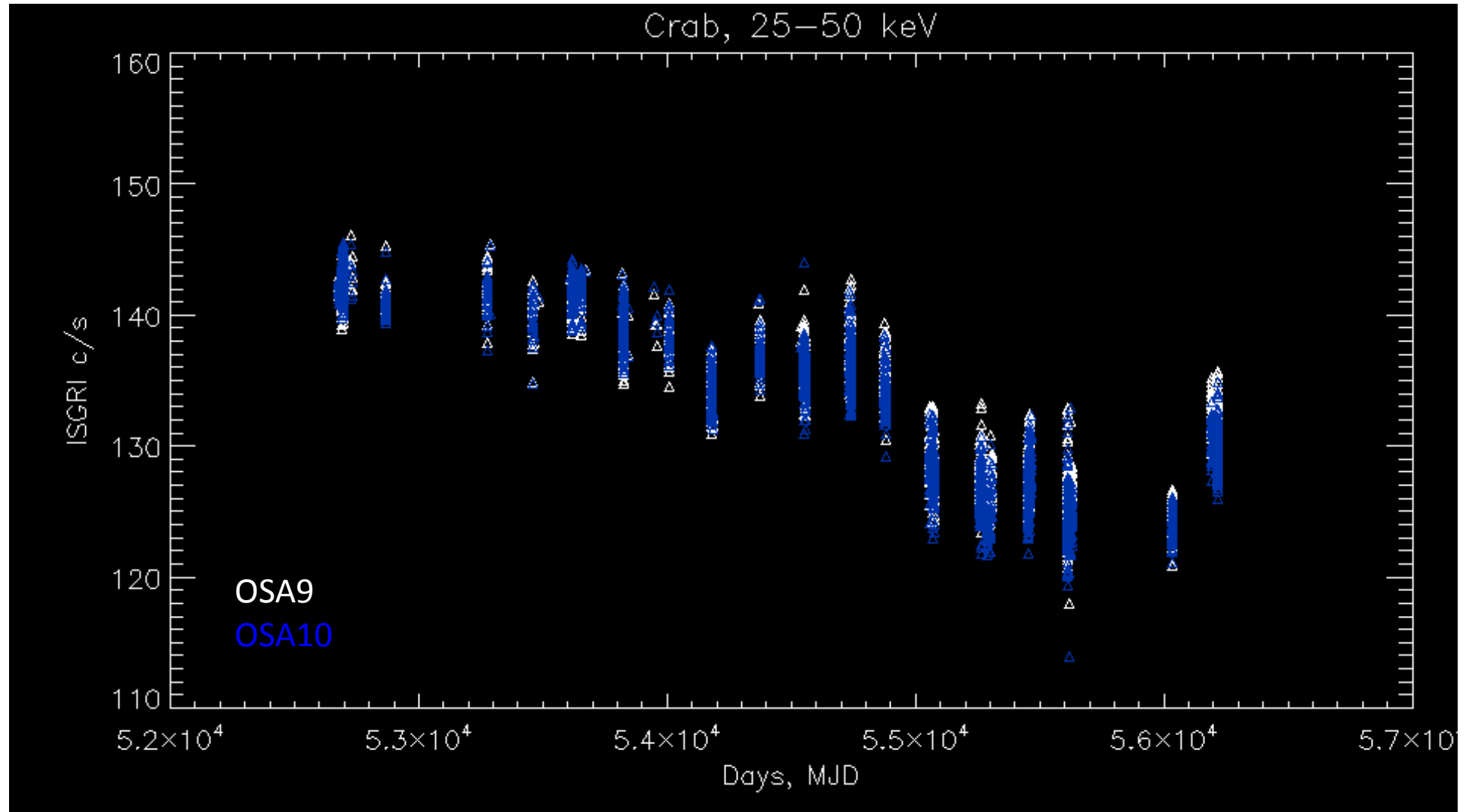
$$\Gamma_1 = 2.08 \pm 0.02,$$
$$\Gamma_2 = 2.28 \pm 0.08$$
$$E_{\text{break}} = 111 \pm 8 \text{ keV}$$

Flux 20-100 keV:
0.274 ph cm⁻² s⁻¹

$$\chi^2 = 28.4/38\text{dof}$$

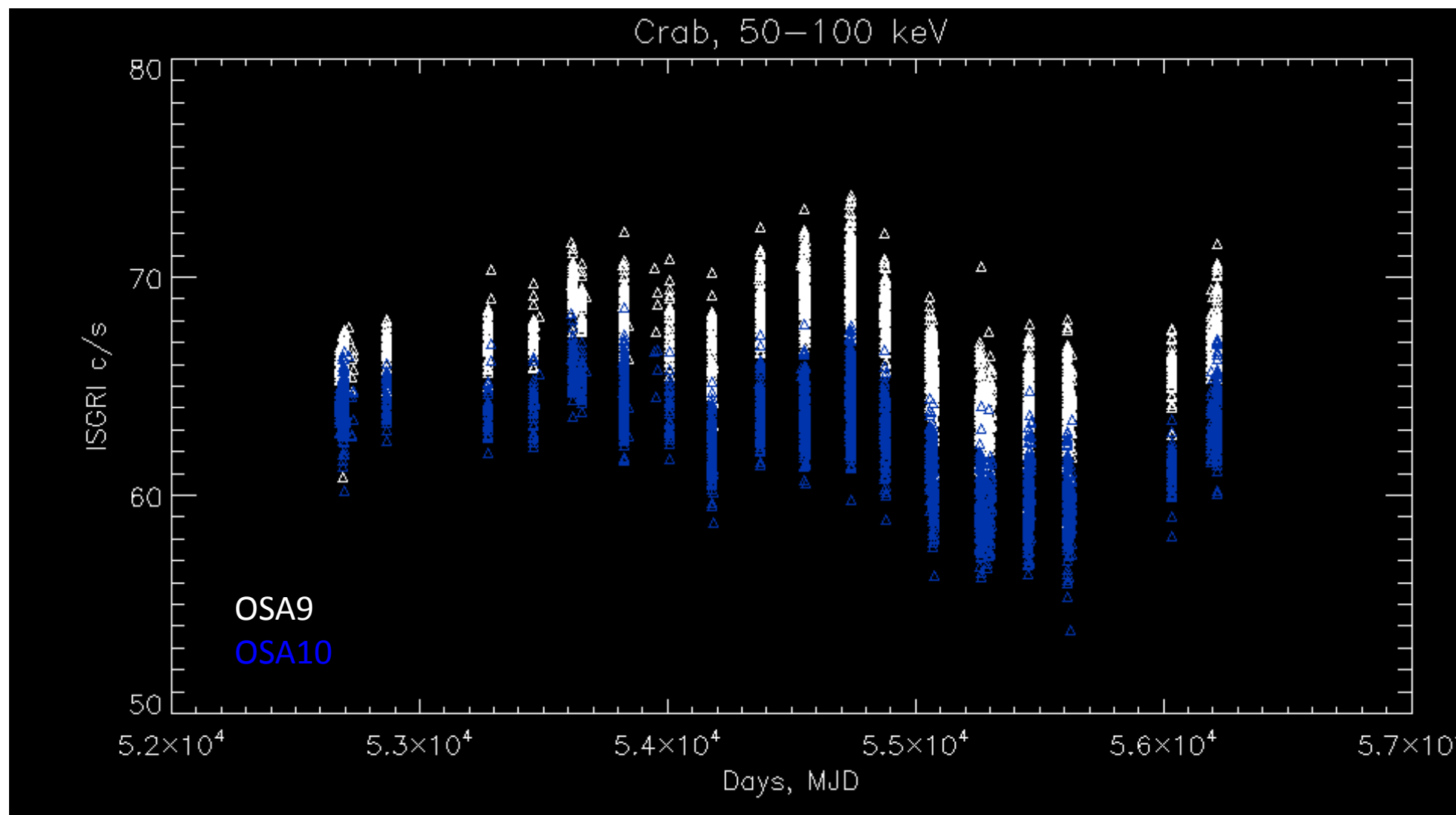
OSA10 vs OSA9

◆ Crab light curve, 10 years, 25-50 keV. No significant difference.



OSA10 vs OSA9 C/s ...not ph!

◆ Crab light curve, 10 years, 50-100 keV. Rates are $\sim 5\%$ lower in OSA10.



Cross-calibration activities

- ◆ Recently launched Nustar has pushed cross-calibration studies and nearly simultaneous campaigns with several satellites including INTEGRAL
- ◆ Nearly simultaneous observations of the Crab with Nustar are now available from INTEGRAL. Work is in progress to exploit cross-calibration.
- ◆ Data from G21.5-0.5 and 3C273 will be also used in the future
- ◆ A special $E > 10$ keV session on cross-calibration will be organized at the 8th IACHEC (2013 March, 25-28 at Hothorp Hall, UK)