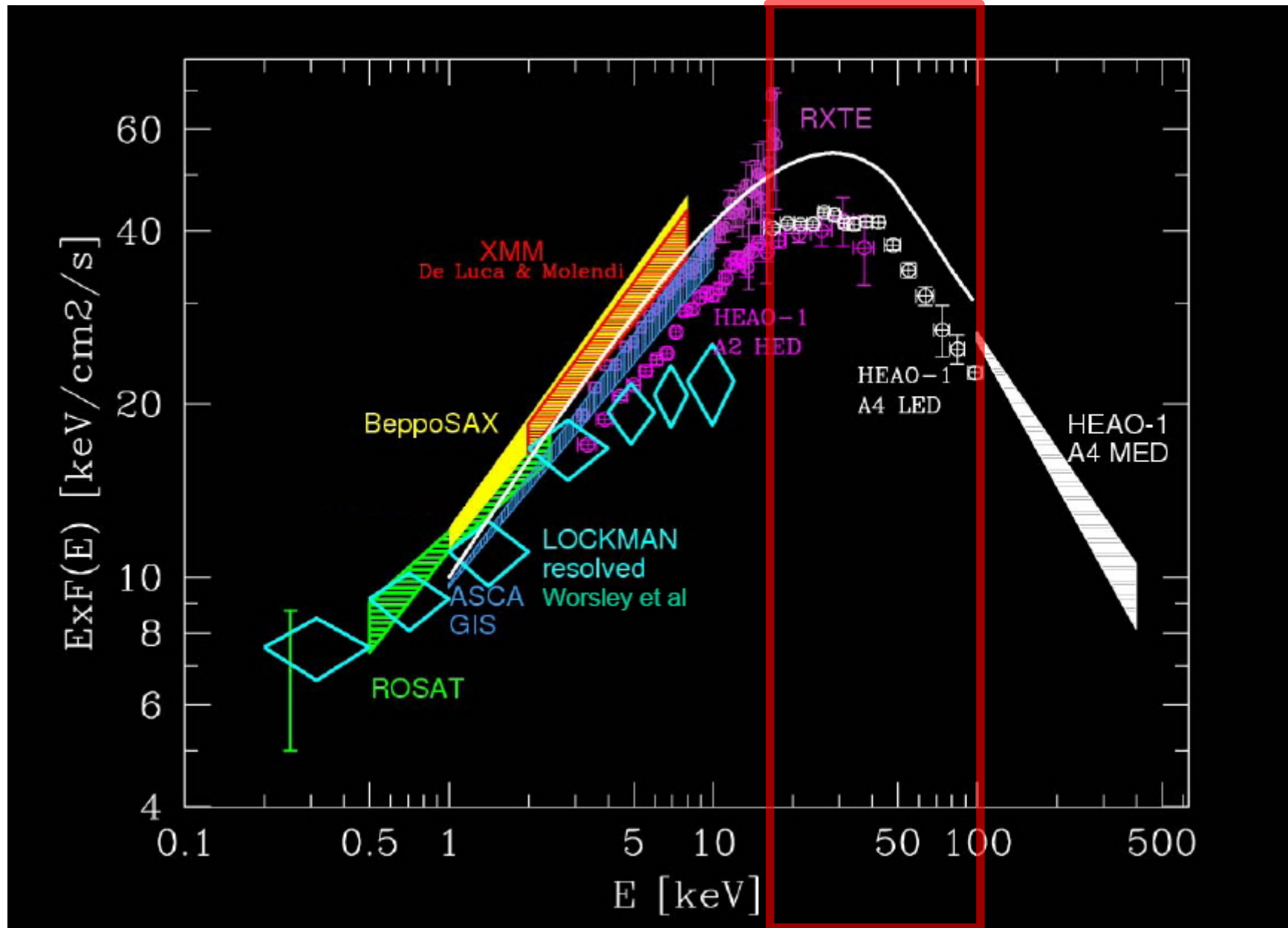
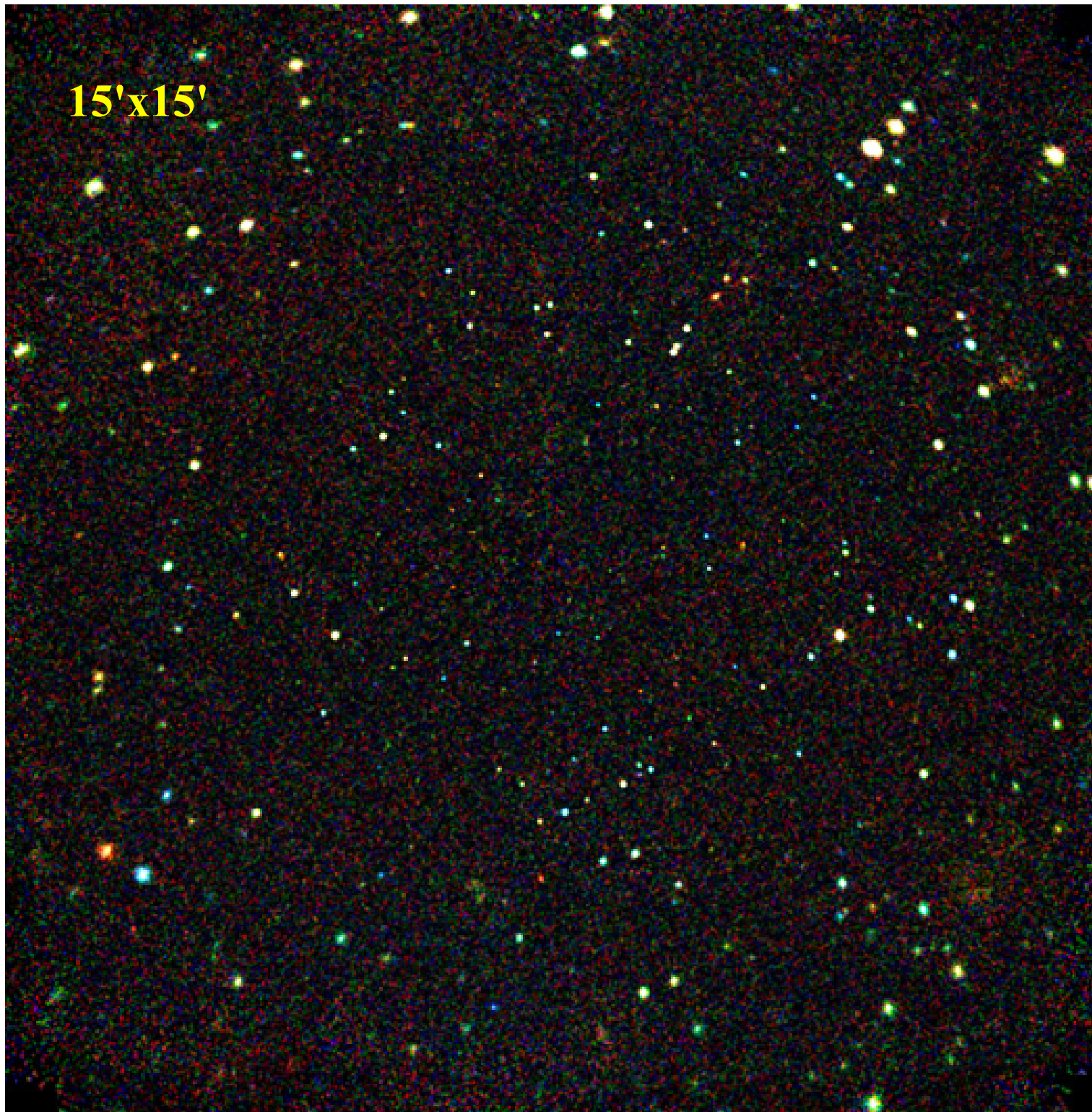


INTEGRAL observations of the Earth



15'x15'

Chandra
0.5-2 keV
80-90% resolved



45x45 deg

XSS/IGR J12389-1614

INTEGRAL
20-60 keV
3% resolved

ESO 511-G030

IC 4329A

MCG-6-30-15

ESO 323-G077

NGC 4507

Cen A

IC 4518A

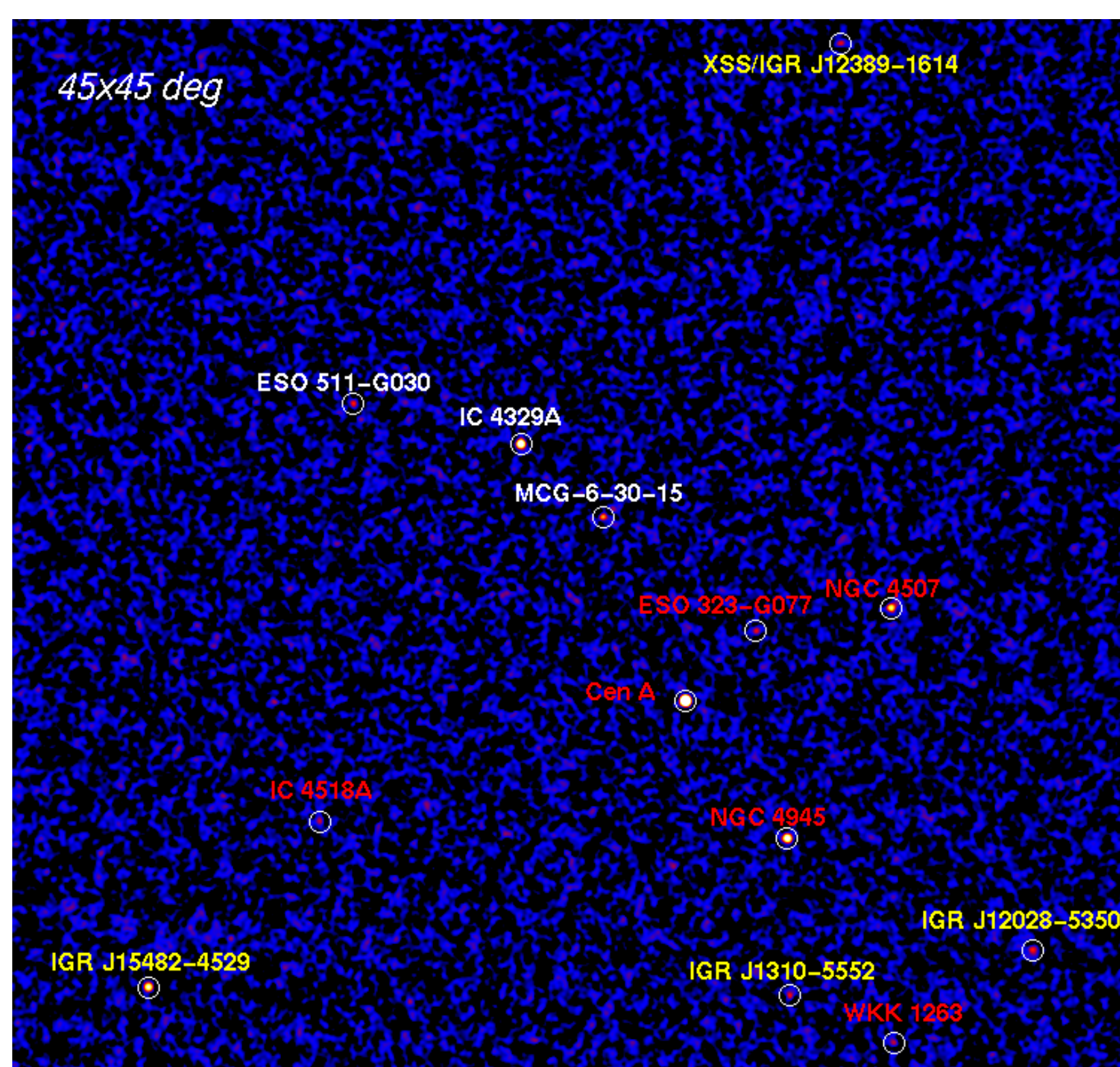
NGC 4945

IGR J12028-5350

IGR J15482-4529

IGR J1310-5552

WKK 1263



45x45 deg

XSS/IGR J12389-1614

INTEGRAL
20-60 keV
3% resolved

ESO 511-G030

ESO 323-G077

NGC 4507

Cen A

IC 4518A

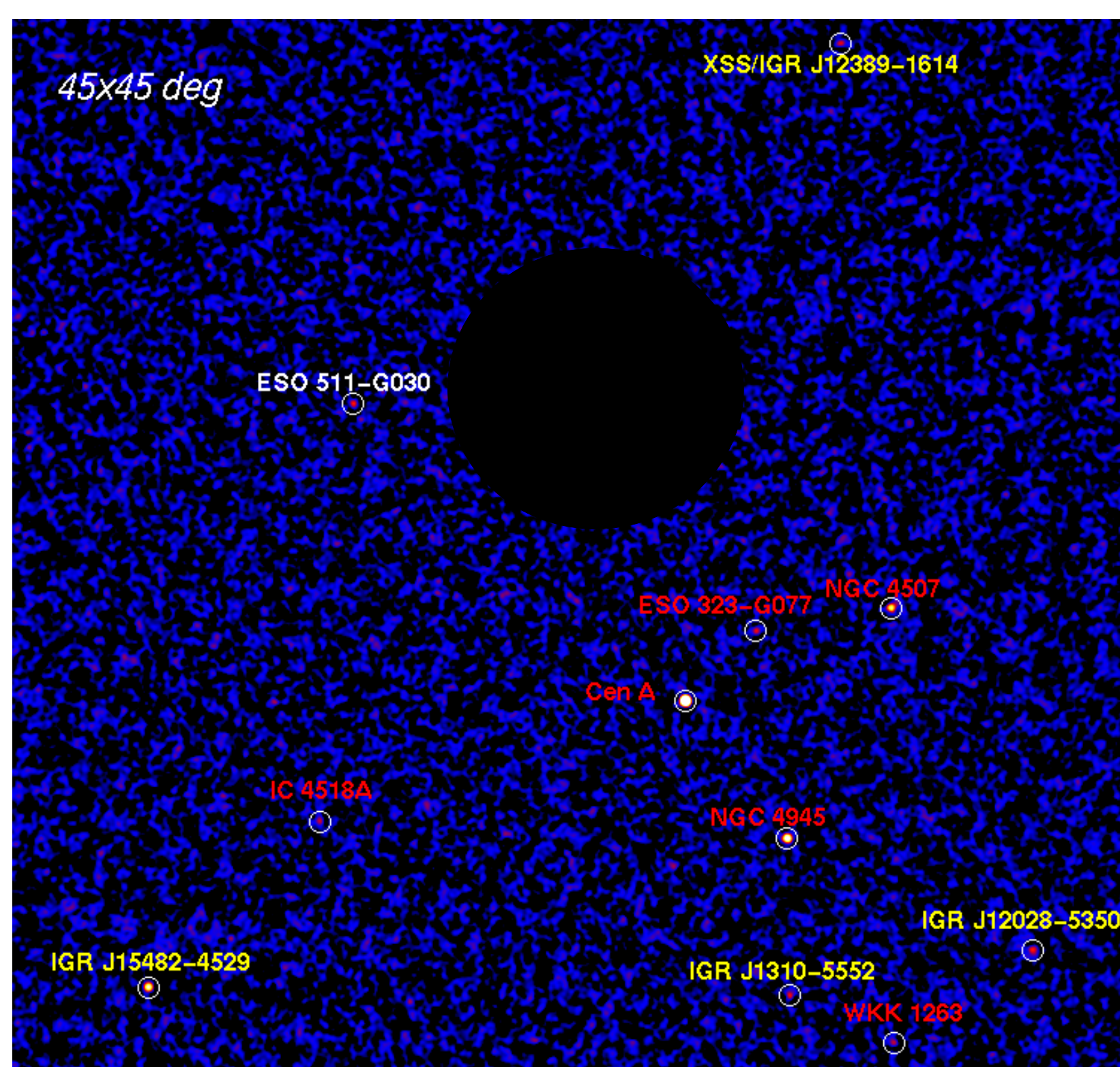
NGC 4945

IGR J12028-5350

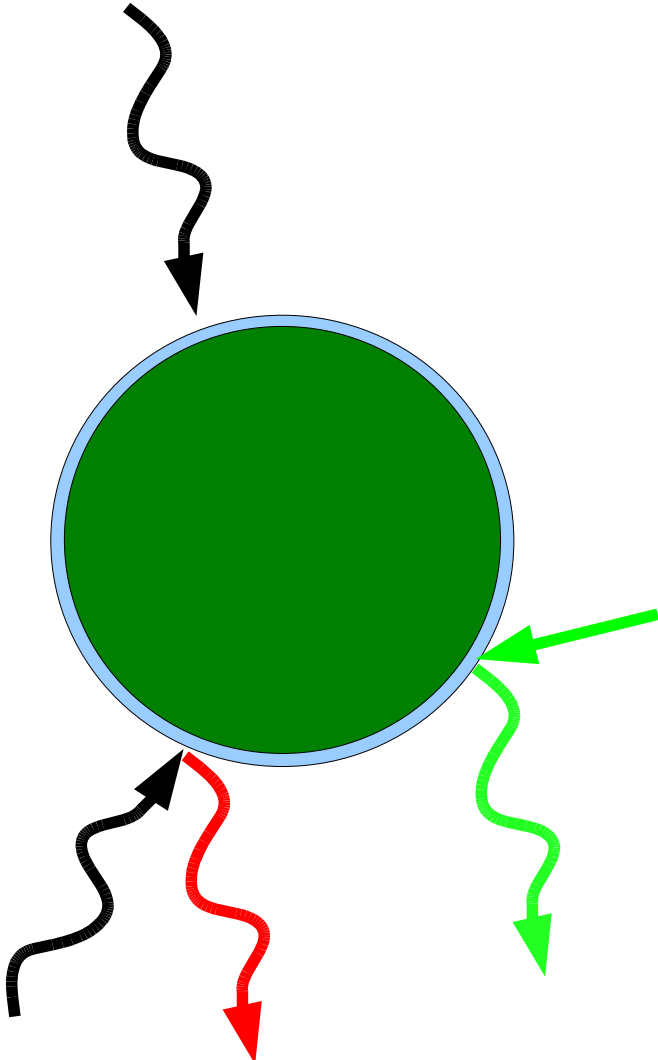
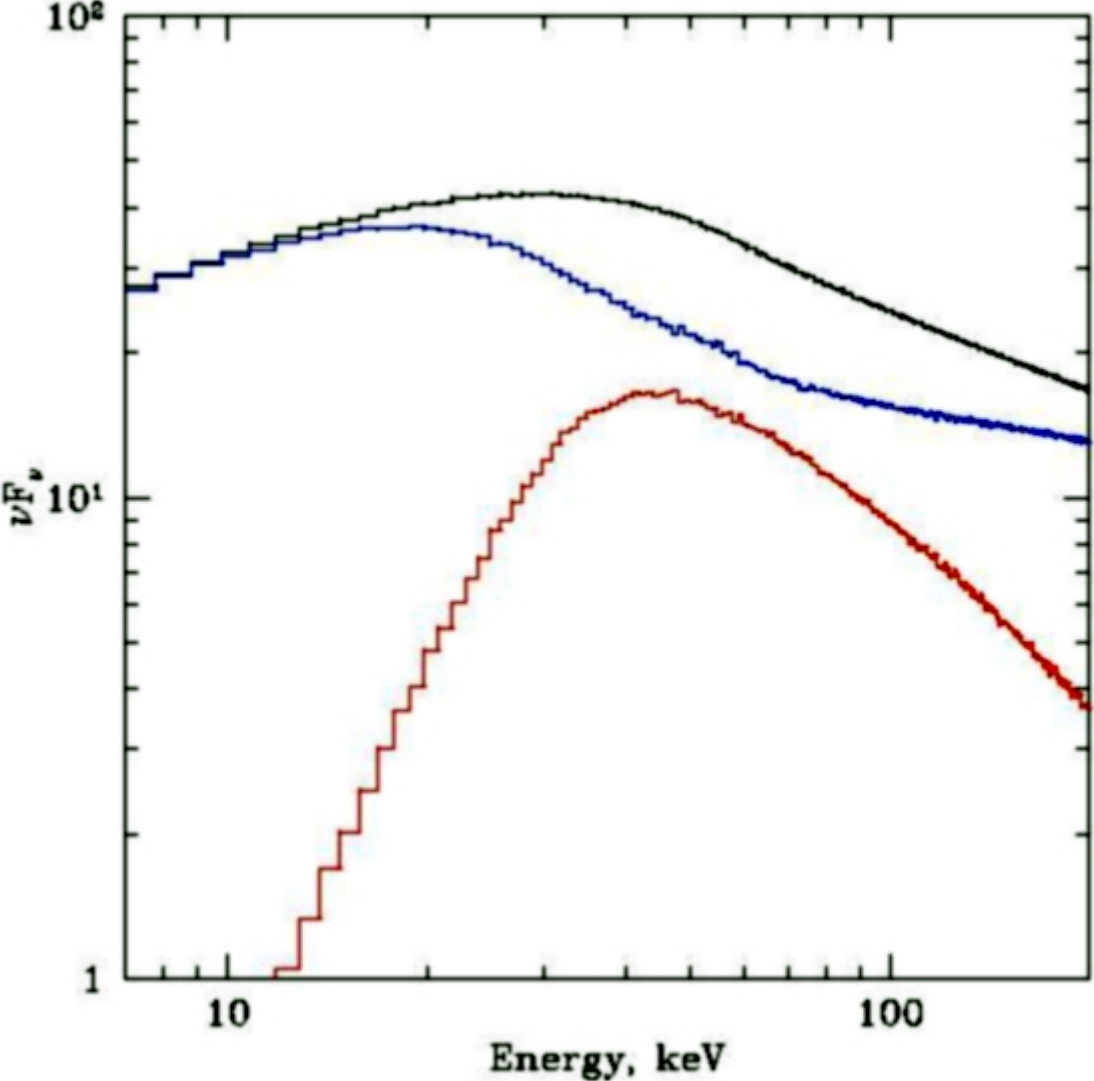
IGR J15482-4529

IGR J1310-5552

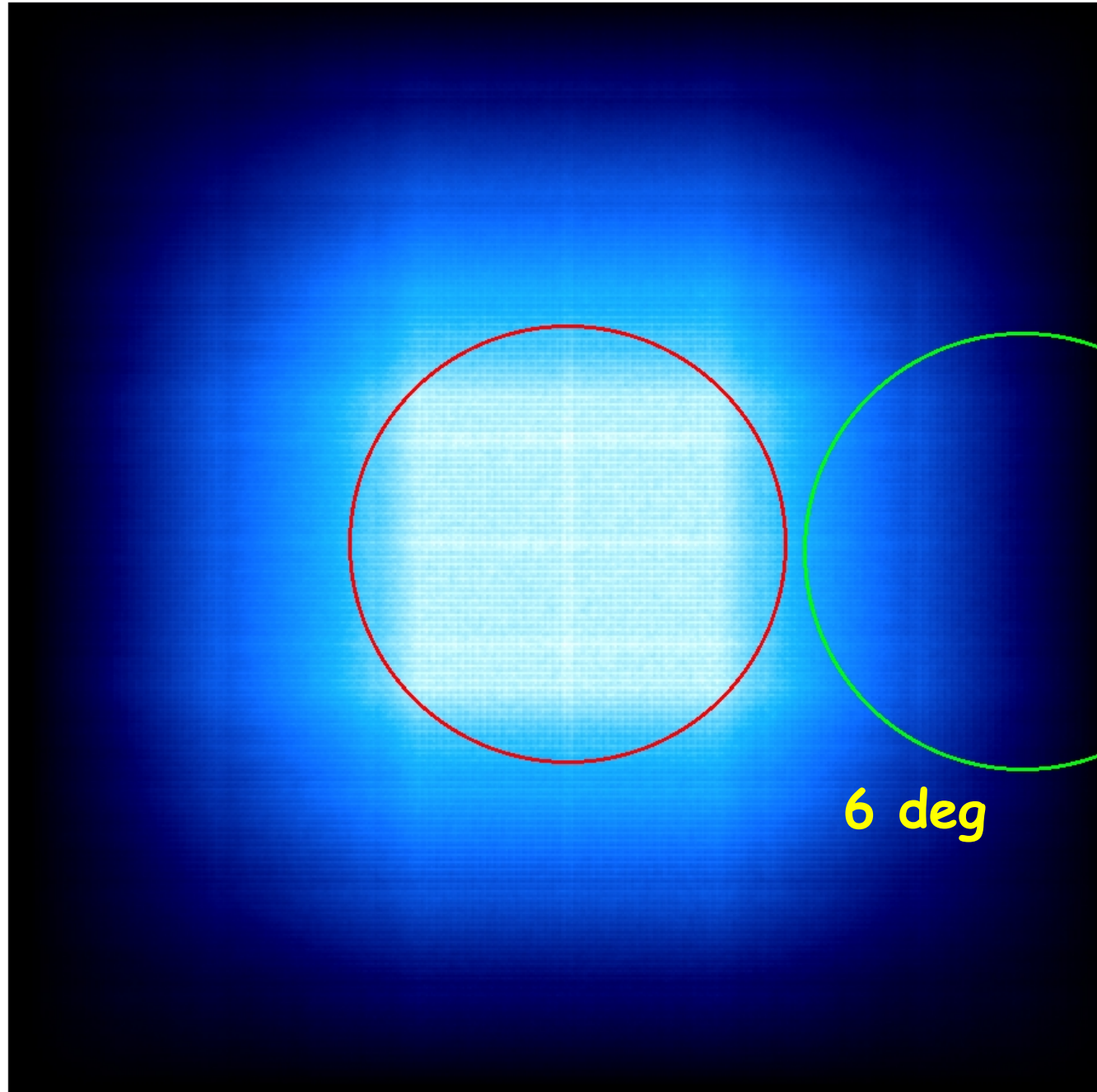
WKK 1263

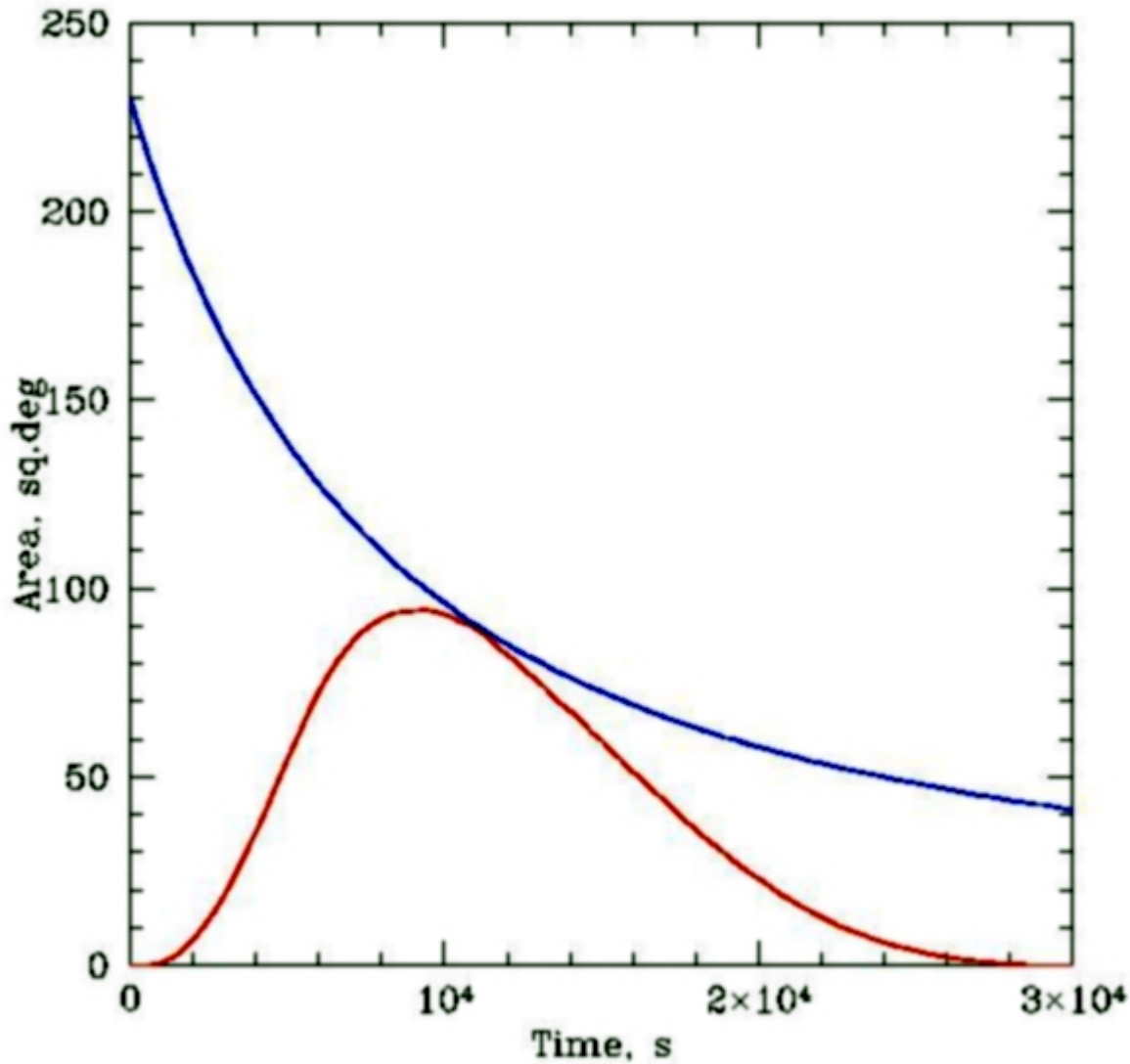


X-ray albedo for N+O atmosphere

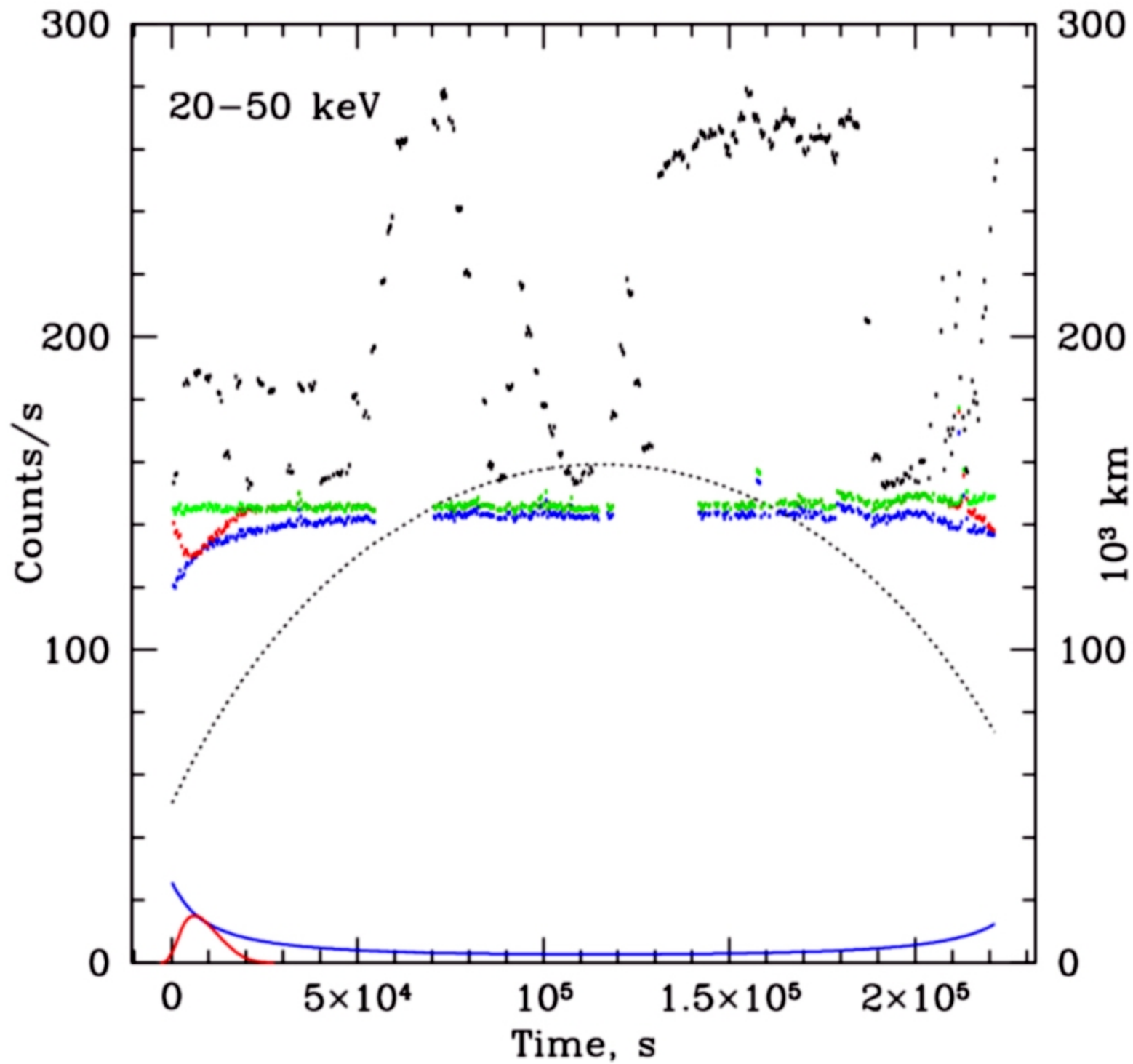


30 degrees



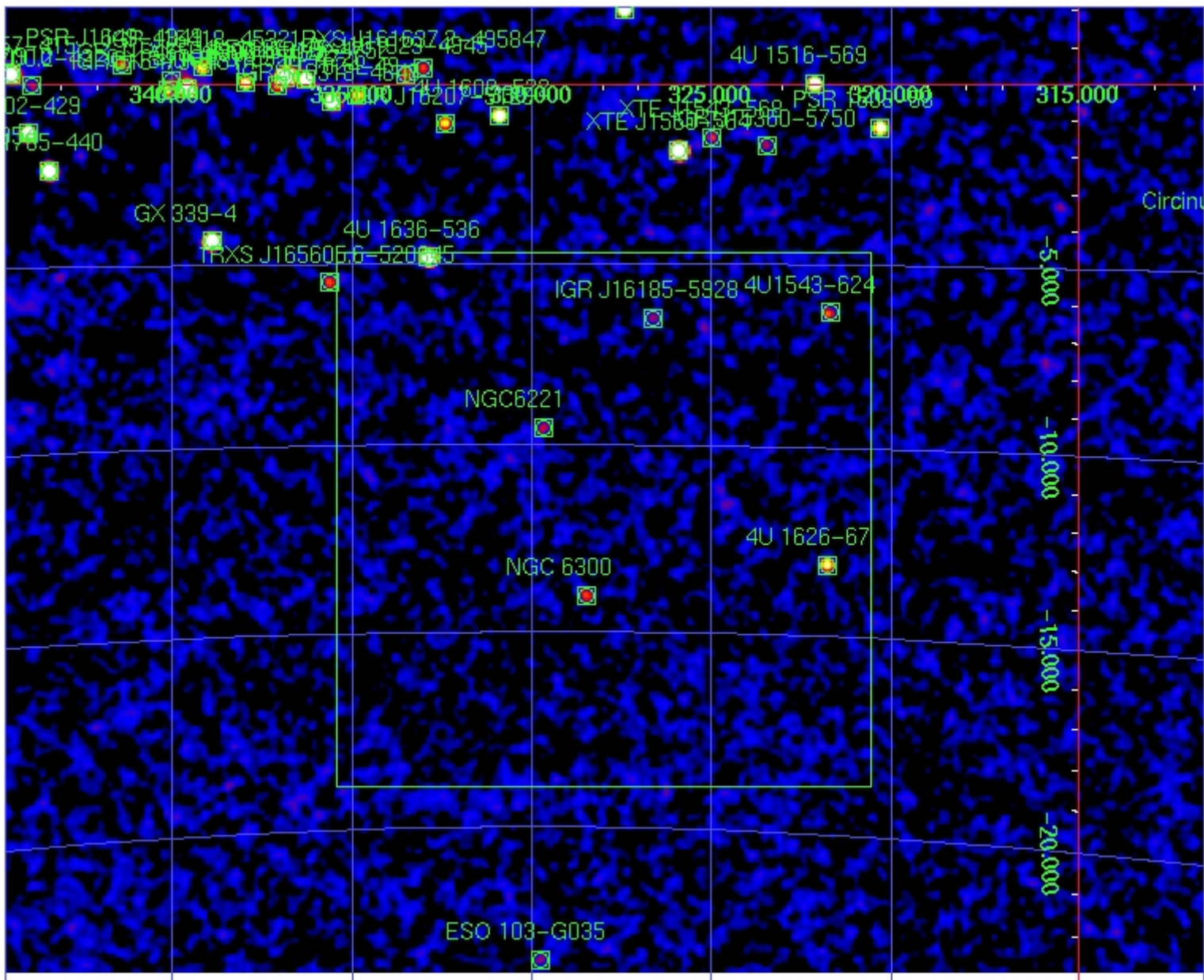


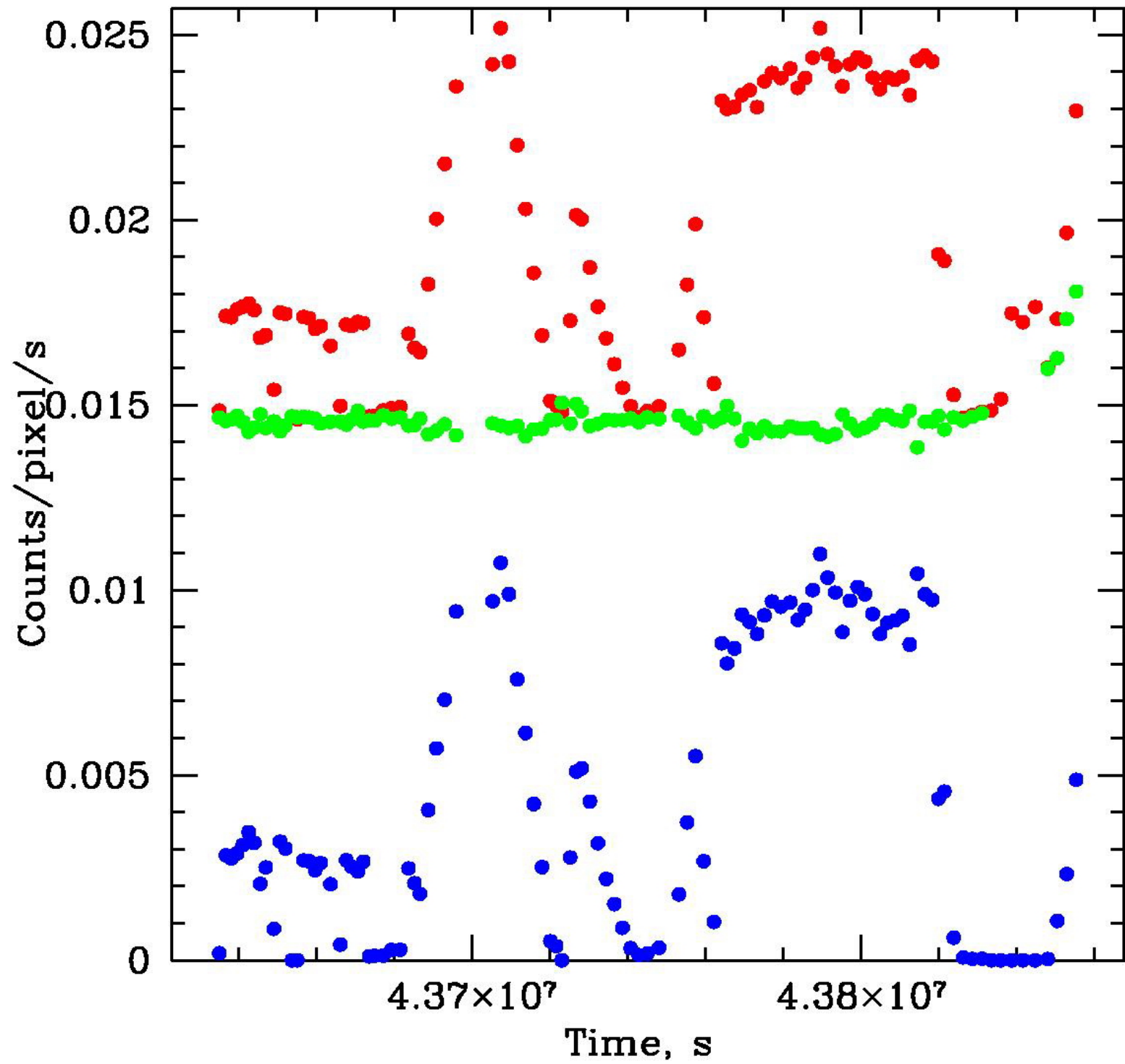
**30 ksec observation
equivalent to:
8 ks of on-Earth pointing
@ 60000 km**

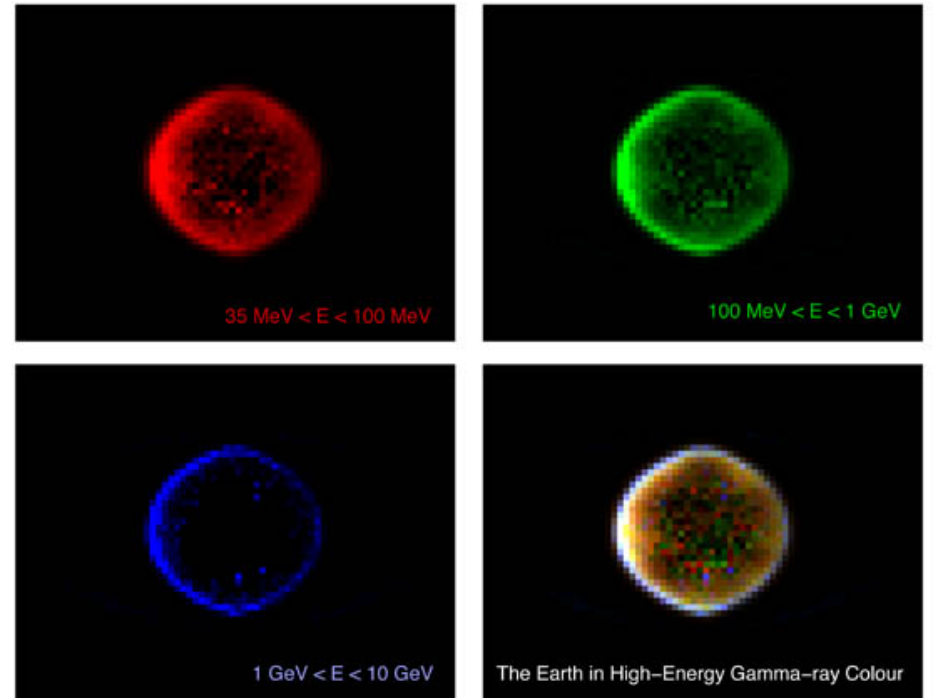
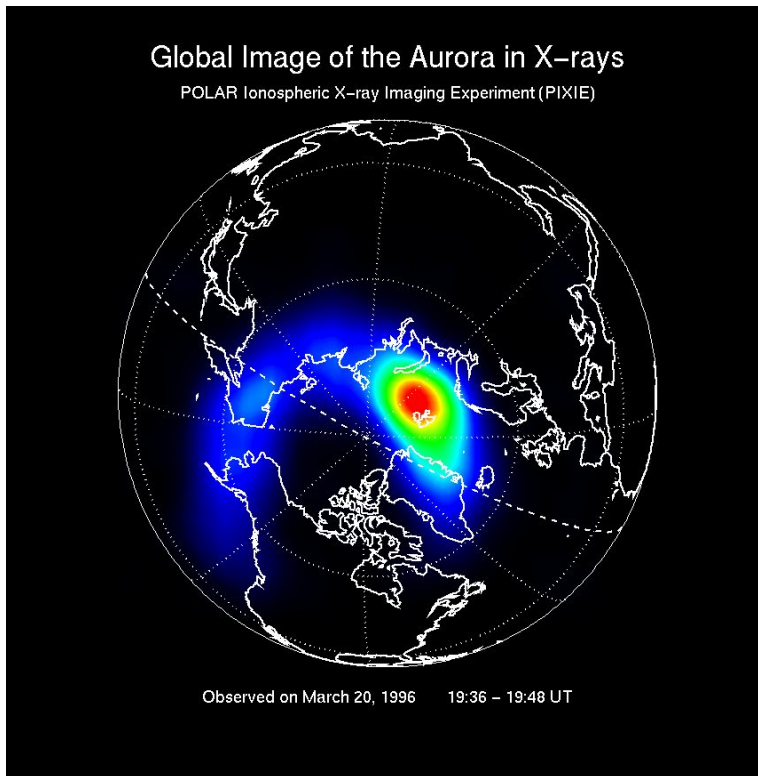


S/N (with X-ray albedo, but without atmospheric emission)

E1	E2	S/N (8 ksec)	S/N (50 ksec)
17	22	70	173
22	29	63	156
29	37	49	122
37	49	37	92
49	63	21	53
63	82	12	30
82	106	10	25
106	138	7	18
138	180	2.8	7

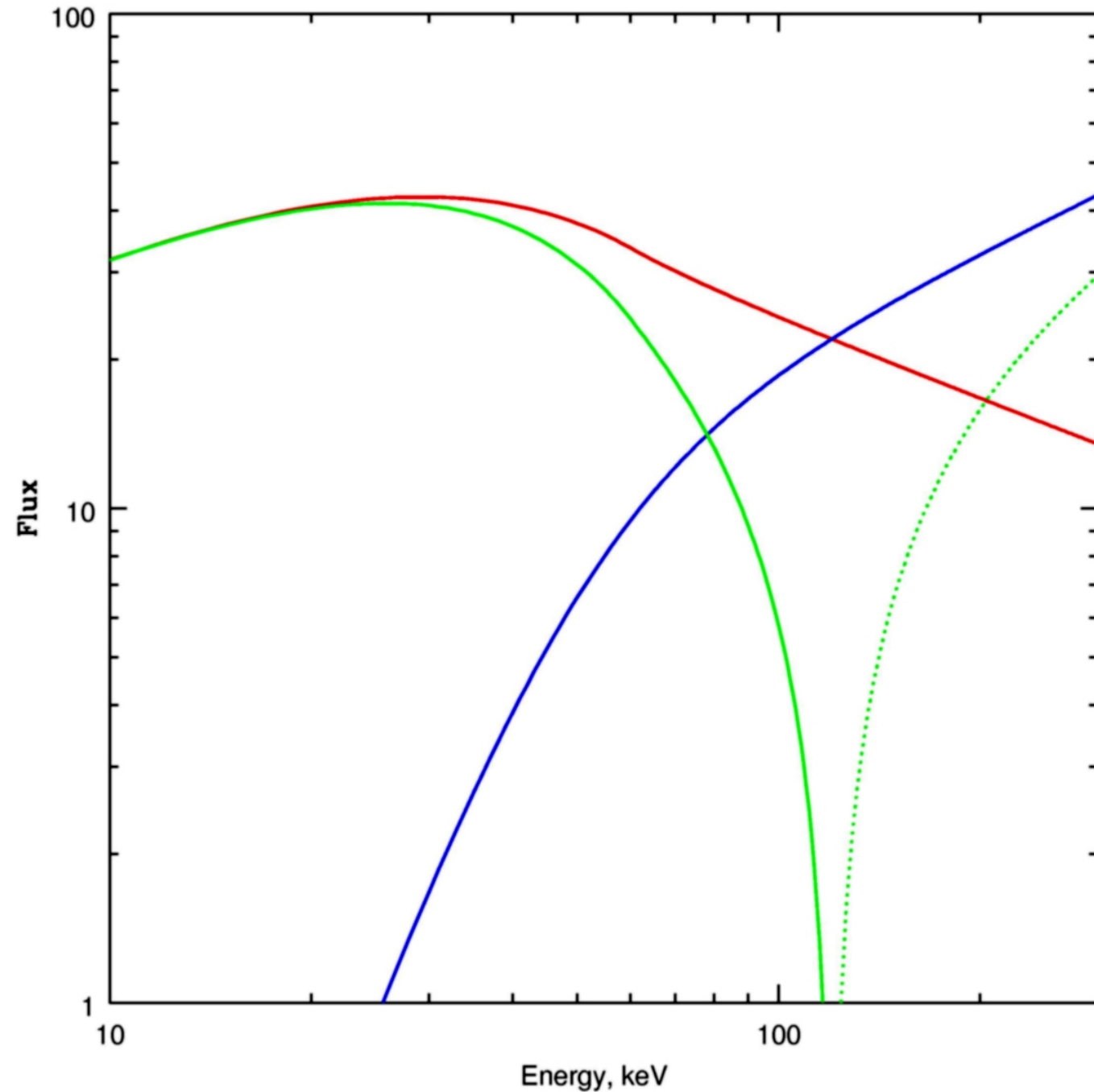






Data: Petry, UMBC & NASA/GSFC, 2004

$E < 100 \text{ keV}$: negative CXB
 $E > 100 \text{ keV}$: positive (atm.emission)



CXB: dark disk

**Atmospheric emission:
F(Rigidity, r/R, Day/Night)**

**Pseudo-imaging at
energies > 50 keV;
errors increase by a
factor 3-5, depending
on the model**

511 keV line: $\sim 4 \cdot 10^{-3} \text{ phot/s/cm}^2$

E1	E2	S/N (8 ksec)	S/N (50 ksec)
17	22	70	173
22	29	63	156
29	37	49	122
37	49	37	92
49	63	21	53
63	82	12	30
82	106	10	25
106	138	7	18
138	180	2.8	7

Divide by 3-5

- 50 ks at the beginning of revolution
- “Empty field” as the background (behind the Earth)
- Any orientation with respect to the Sun
- Orientation is needed a posteriori
- JEM-X+ISGRI + SPI (~100 mCrab @ 100 keV @ 50000 km
~ 50 mCrab @ 500 keV
~10–20 mCrab @ 2 MeV)
- Above ~100 keV positive source

Earth Observation attitudes (at 6 hrs after perigee exit)
Start rev = 384; end rev = 818

